



# Improving Gravitational Wave Searches by Using Multivariate Classifiers to Incorporate Auxiliary Channel Information

Kari Hodge

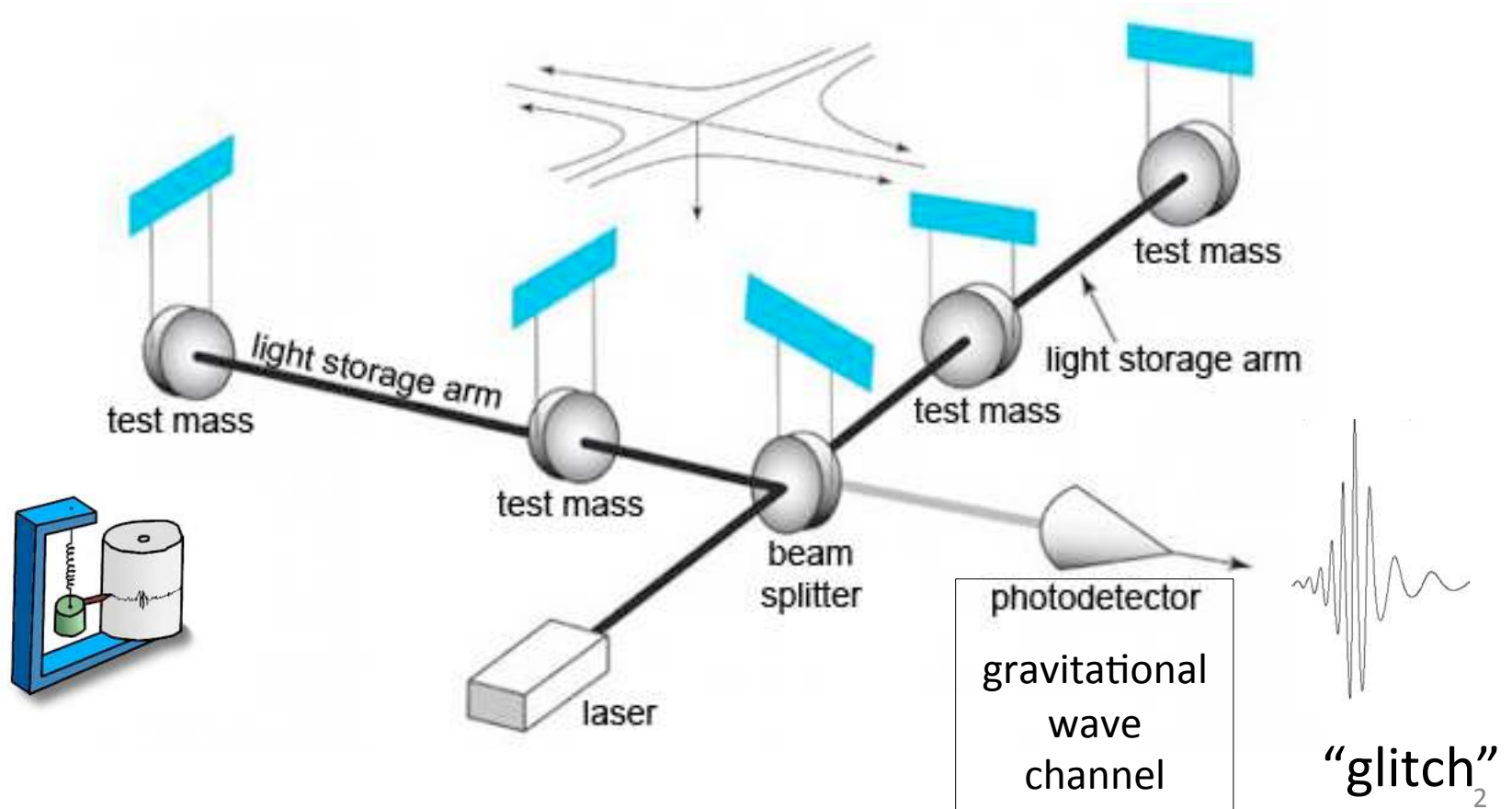
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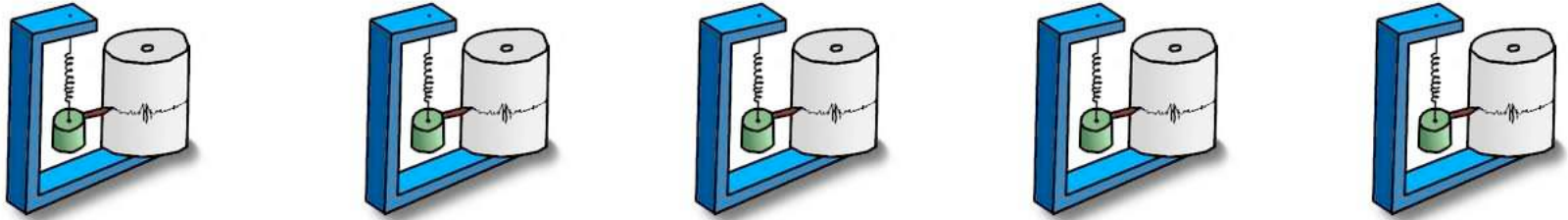


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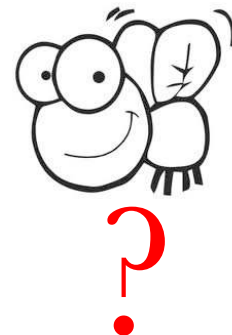
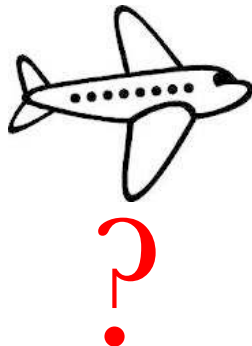
# Laser Interferometer Gravitational-Wave Observatory (LIGO)





The detectors have hundreds of such **auxiliary channels** monitoring the instrument and its environment

In practice, a multitude of them witness “something” and the cause of the glitch is not clear



# Auxiliary Channels

- Environmental:
  - seismometers
  - microphones
  - magnetometers
- Instrumental:
  - beam splitter motion
  - mirrors' angular alignment
  - common arm

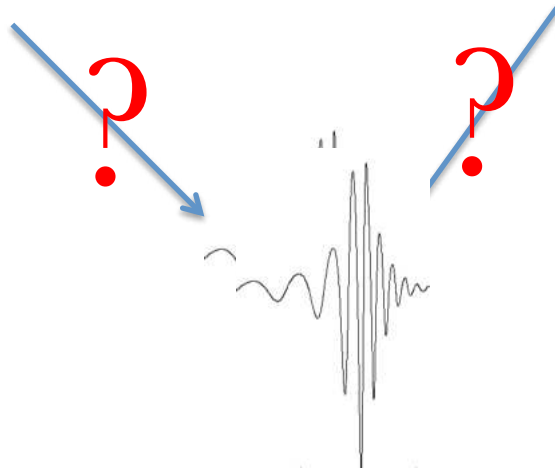
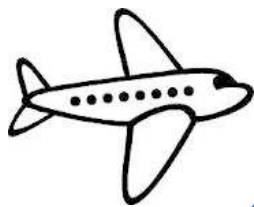


# Problem:

Glitches increase false alarm rates, obscuring detections and raising upper limits

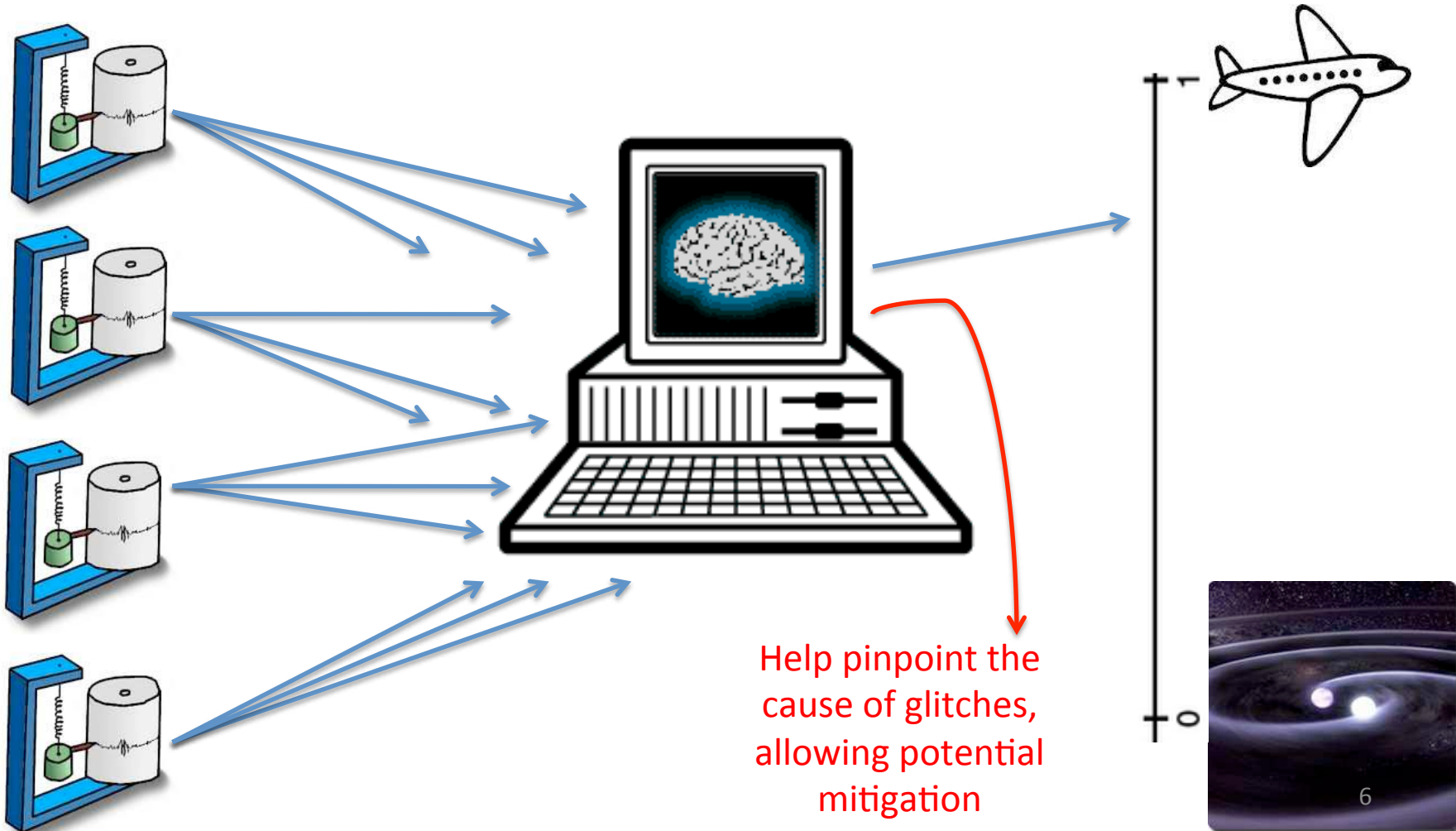
So we'd like to be able to either:

- veto these times prior to analysis  
(historical method, based on at most a few channels per veto)
- re-rank events based on the result of auxiliary channel analysis  
(difficult analytically because of the sizable number of channels)



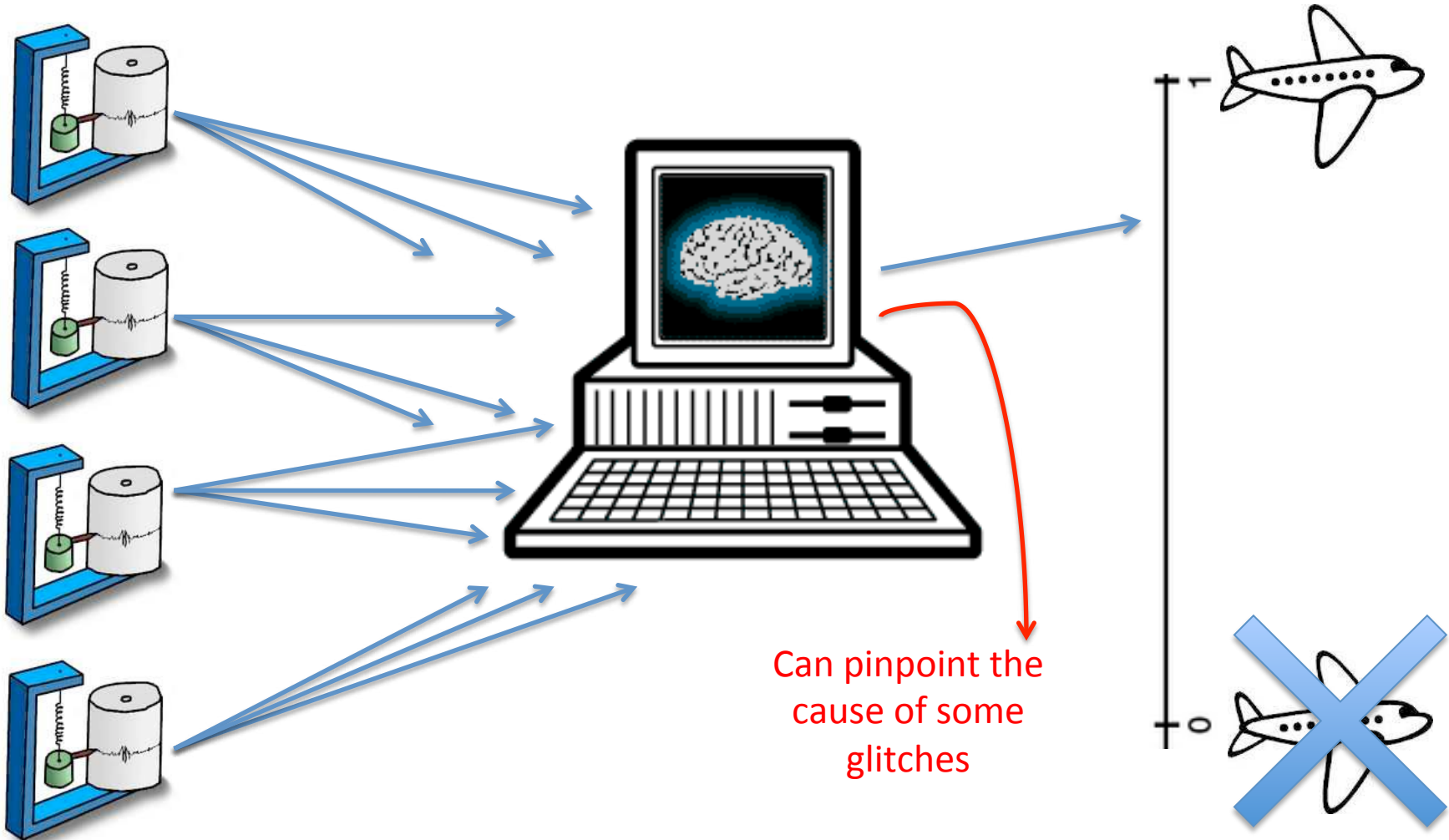
# Ideal Solution:

Combine information from these auxiliary channels into a Multivariate Classifier (MVC) with existing search algorithms, let it determine how likely it was that what was seen in the GW channel was a glitch of environmental/instrumental origin as opposed to an actual gravitational wave



# So Far:

Completed half of this – separating generic artifacts from “clean” times.  
Future work will couple this to specific searches.







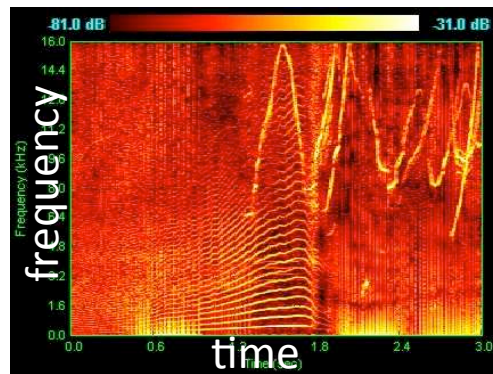
# Input to the MVCs

- classifiers are constructed with a training set – events with known class
- **IMPORTANT:** the training set must be unbiased
  - can't contain any information from the gravitational wave channel
  - glitch set shouldn't contain gravitational waves



# Input to the MVCs

- Training set:
  - Class 1: the glitches (what we are looking for)
    - identified by a Kleine-Welle trigger in the gravitational-wave channel above a nominal threshold
  - Class 0: “clean” times
    - at least 100 ms away from a Class 1 sample



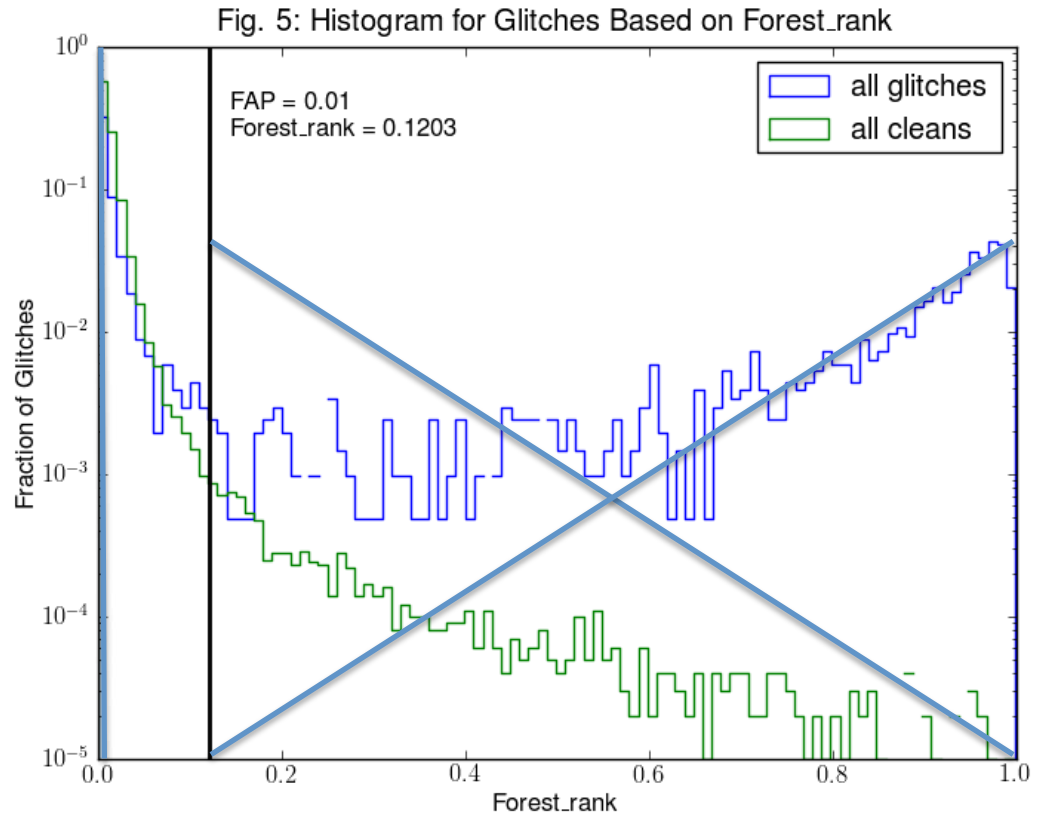
Kleine-Welle captures excess power in the time-frequency plane via wavelet analysis

figure: generic spectrogram

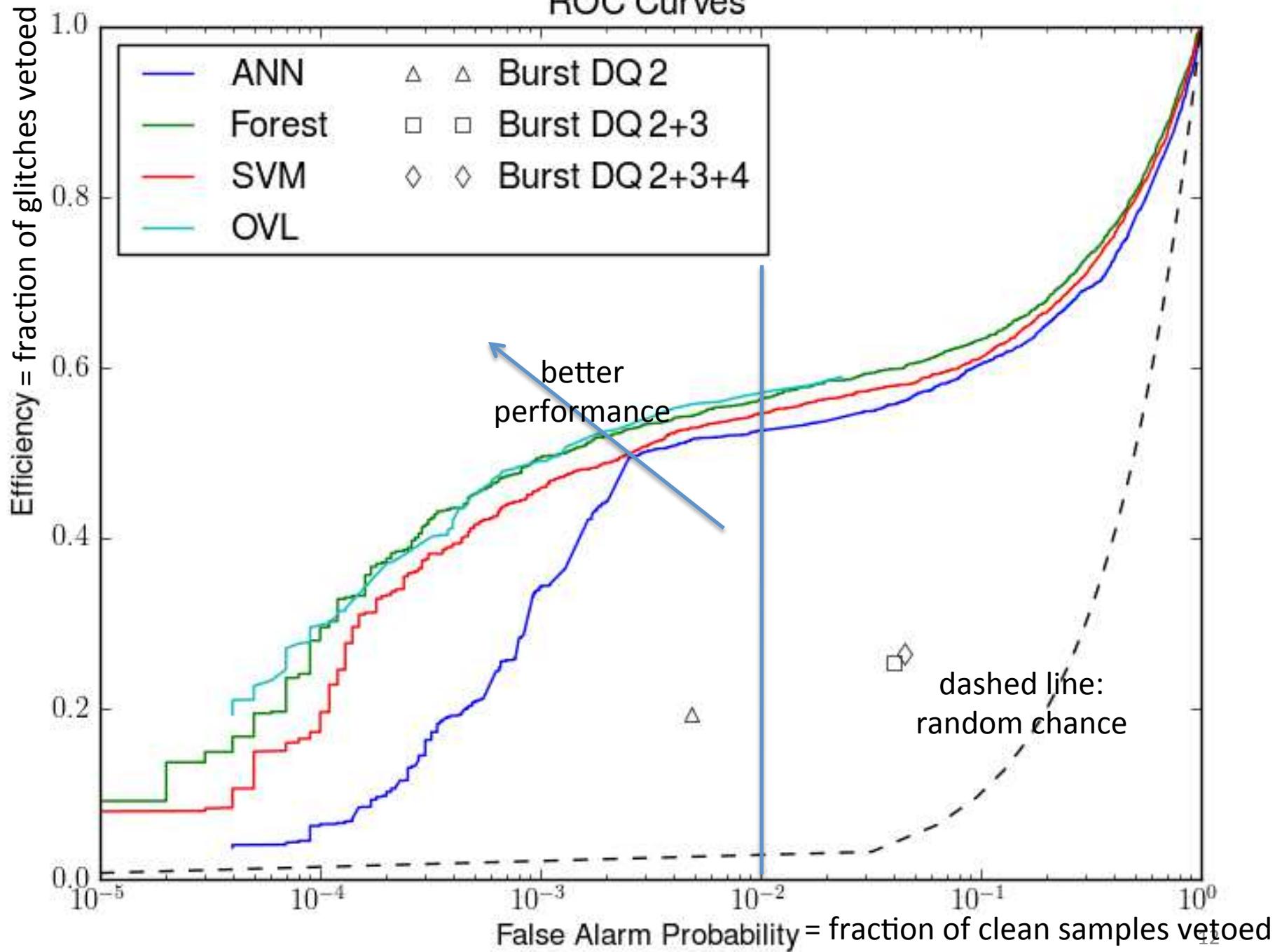
Each event is parameterized by 5 numbers per auxiliary channel. These numbers are extracted from the auxiliary channels with the same Kleine-Welle wavelet analysis.

# Output of the MVCs

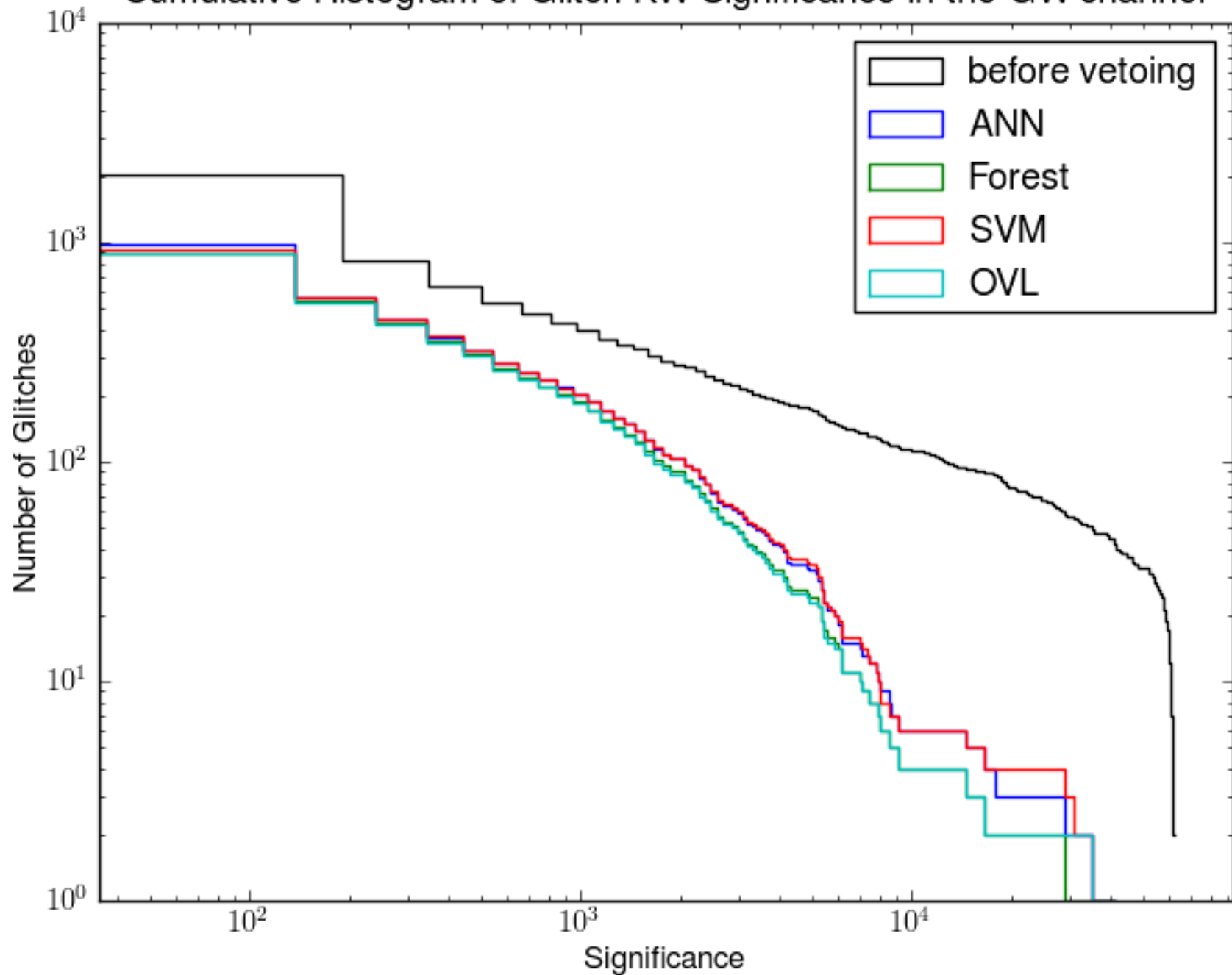
- Once the MVC is trained, you can feed it an event and it will spit out a number between 0 and 1, where 1 indicates it is more glitch-like.
- By picking a threshold on this continuous rank, one can veto any GPS time where the rank exceeds this threshold



# ROC Curves

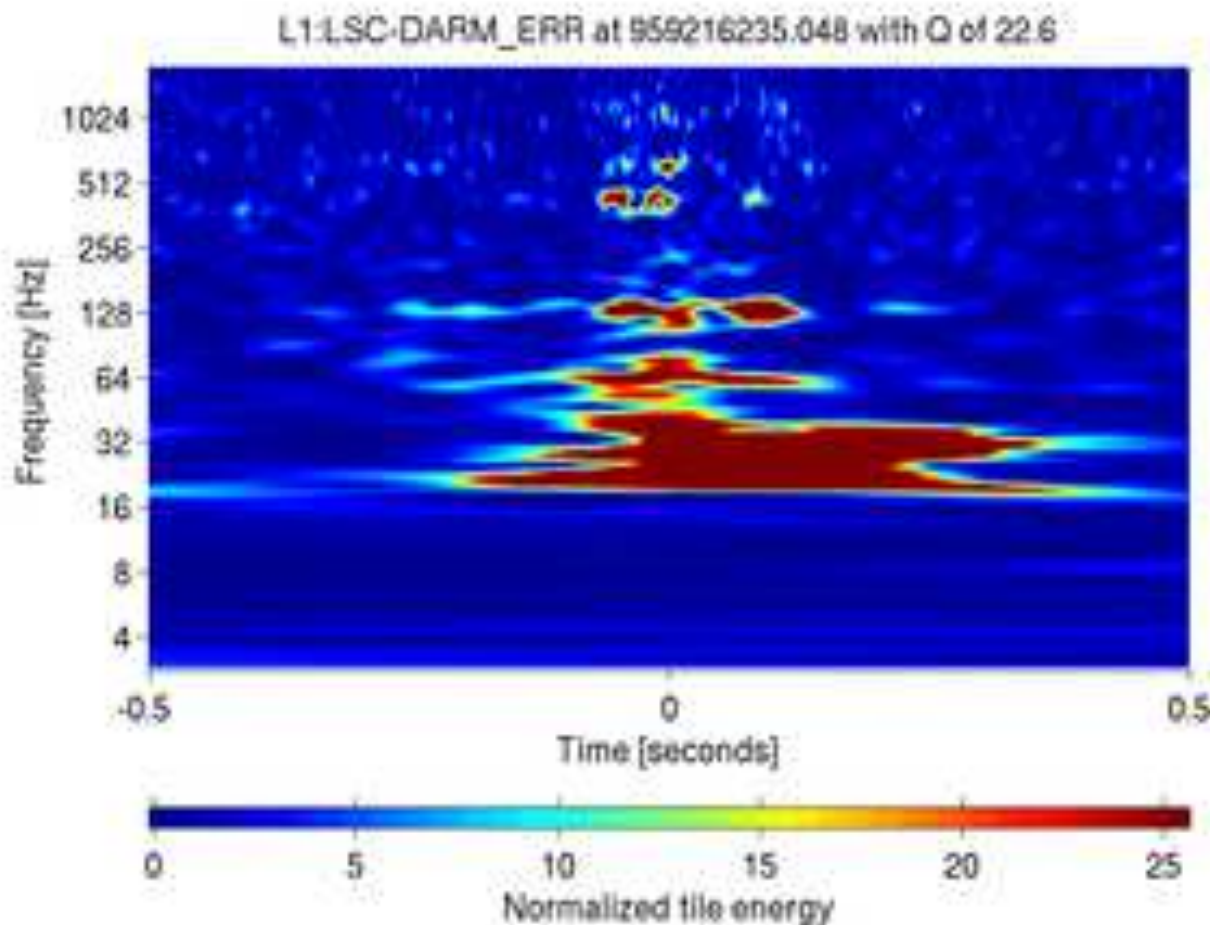


Cumulative Histogram of Glitch KW Significance in the GW channel



# Detector Characterization

- We can identify the channels witness to specific classes of glitches and feed this information back to instrumentalists



Train only on events that have this character in GW channel

The trained forest has the most splits on a specific set of channels

This “jitter” is caused by a non-linear coupling of slow motion of the optics with noise at higher frequencies in the output mode cleaner

# Conclusions

- The key to distinguishing a gravitational wave from an instrumental or environmental glitch is in the auxiliary channels
- Once trained, the MVCs can classify new events rather quickly
- The different classifiers find mostly the same glitches
- MVCs are not just a black box!

# Future Work

- expansion of input variables will lead to better performance
  - include other time-frequency analyses (e.g. Omega) to capture more information
  - direct channel readout for slow channels
- classification of different glitch types
- diagnosis of the physical causes of more classes of glitches (upconversion noise)
- thorough check of veto safety (hardware injections)
- interfacing with CBC and Burst searches
  - define a detection statistic for a search which will include MVC output