

TIGER

A data analysis pipeline for testing general relativity
using compact binary coalescence

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TIGER in short



- Test validity of GR using Bayesian model selection
 - Based on hypotheses instead of values of parameters
 - Make use of full information content
- Tailored to Advanced LIGO/Virgo
 - Suitable for low SNR
 - Moderate number of sources
- Quantify our belief in the validity of GR
 - Even in the presence of spurious effects (e.g. noise)
- Flexible in use
 - Independent of waveform family
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Competing hypotheses: \mathcal{H}_{GR}



- \mathcal{H}_{GR} : Waveform has the form as predicted by GR
- E.g. Post-Newtonian (PN), effective one body (EOB)

Example: Testing the PN phase [1, 2]

- Phase coefficients, ψ_n , predicted by GR and depend on system's characteristics (masses, spins, ...)

$$\Psi(v) = \sum_n \left[\psi_n + \psi_n^{(l)} \log\left(\frac{v}{c}\right) \right] \left(\frac{v}{c}\right)^{n-5}$$

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- $\mathcal{H}_{\text{modGR}}$: One or more terms in the waveform is not as predicted by GR

Example: Testing the PN phase [1, 2]

- One or more ψ_n not as predicted by GR
- Split $\mathcal{H}_{\text{modGR}}$ into testable sub-hypotheses

$$\mathcal{H}_{\text{modGR}} = \bigvee_{i_1 < i_2 < \dots < i_k} H_{i_1 i_2 \dots i_k} \quad (1)$$

- Define $H_{i_1 i_2 \dots i_k}$: $\psi_{i_1}, \dots, \psi_{i_k}$ do not have the functional dependence on the masses as predicted by GR, but all other $\psi_j, j \notin \{i_1, i_2, \dots, i_k\}$ do have the dependence as in GR

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Odds ratio



- Analyse signal within stretch of data, d
- Construct the odds ratio as our figure of merit

$$O_{\text{GR}}^{\text{modGR}} = \frac{P(\mathcal{H}_{\text{modGR}}|d, I)}{P(\mathcal{H}_{\text{GR}}|d, I)} \quad (2)$$

- $O_{\text{GR}}^{\text{modGR}} > 1$ favours $\mathcal{H}_{\text{modGR}}$, $O_{\text{GR}}^{\text{modGR}} < 1$ favours \mathcal{H}_{GR}
- Naturally allows for the combination of sources

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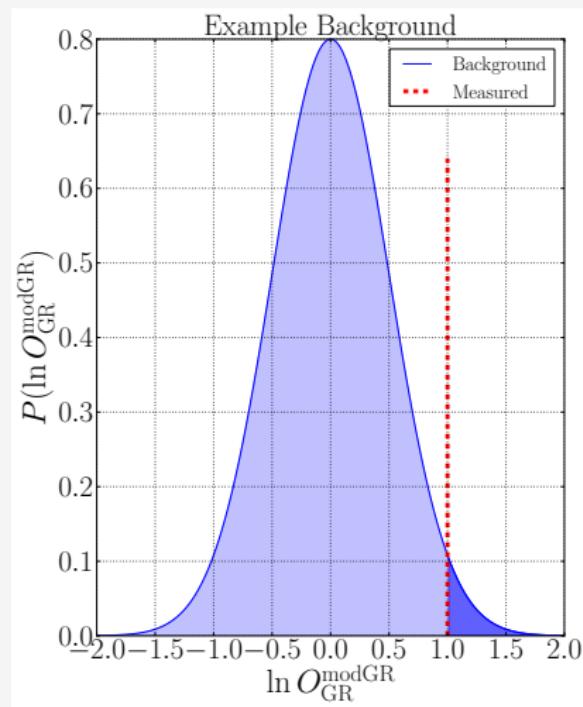
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Constructing the background

- Noise can introduce false positives
- Construct a *background* by analysing GR signals in simulated noise
- Compare measured odds ratio to background to assess the *false alarm probability*

Constructing the background

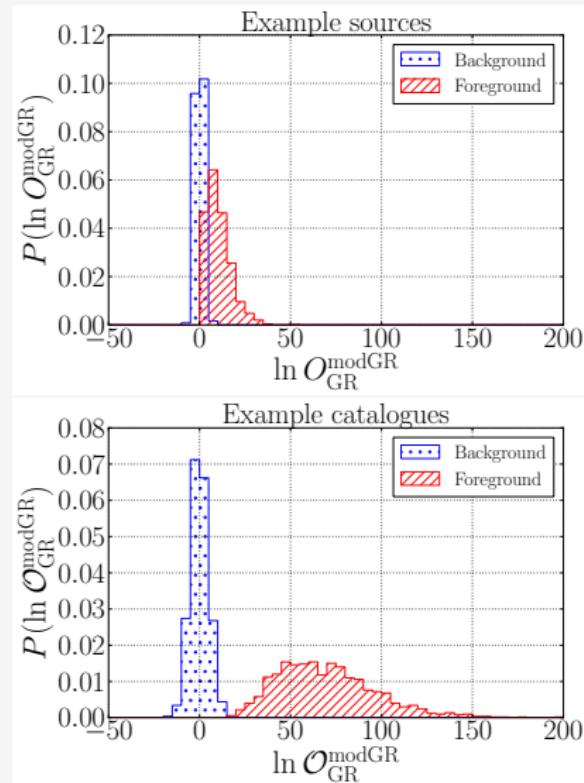
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Testing TIGER (BNS)

Simulated deviations

- Test TIGER by introducing deviations to the waveforms
- Analyse many sources with the same type of deviation, *foreground*
- Compare background to foreground to asses how likely we can distinguish this type of deviation
- Combine sources to increase information



Simulation details



- Advanced LIGO/Virgo network [3, 4]
- BNS systems $M \in [1, 2]M_{\odot}$
- Realistic source distribution [5]
 - $D_L \in [100, 400]\text{Mpc}$
 - $SNR \in [8, 50]$
 - Uniform sky location/polarisation
- Construct background by analysing GR injections
- Construct foreground by analysing deviations

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10% deviation at 1.5PN

- TaylorF2 templates [6]
- Introduce 10% shift at 1.5PN phase term [1]
- Top: single sources, moderate separation
- Bottom: catalogues of 15 sources, complete separation

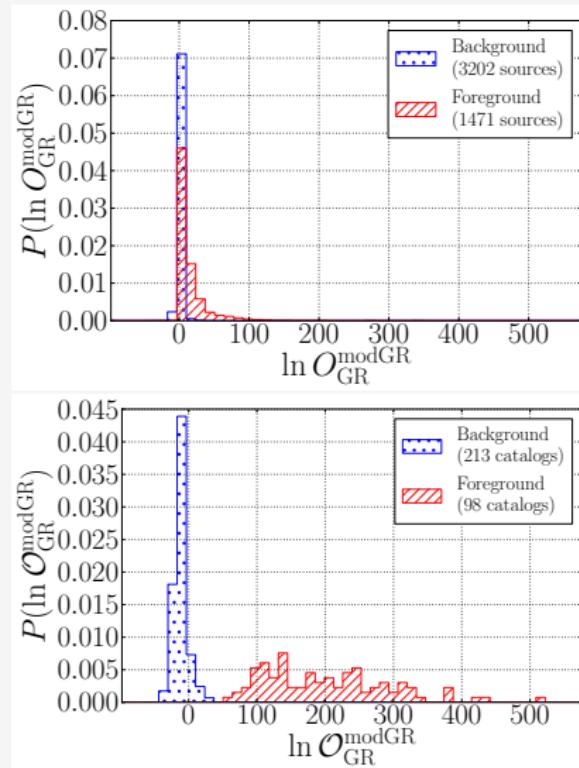
Detect deviations that are included in $\mathcal{H}_{\text{modGR}}$

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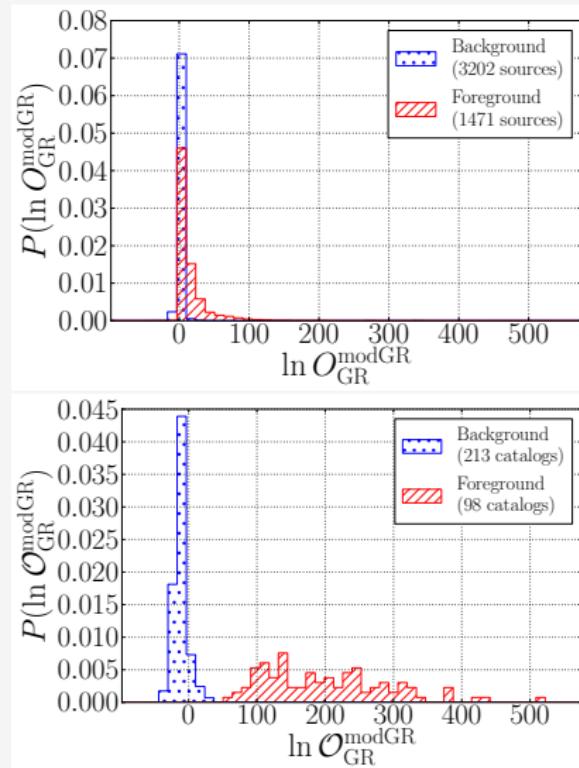


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Mass dependent deviation

- TaylorF2 templates
- Deviation of the form [2]

$$\frac{3}{128\eta} (v/c)^{-6+M/(M_\odot)}$$

- Mass dependent power of velocity
- Fully generic, not included in $\mathcal{H}_{\text{modGR}}$

Sensitive to generic deviations

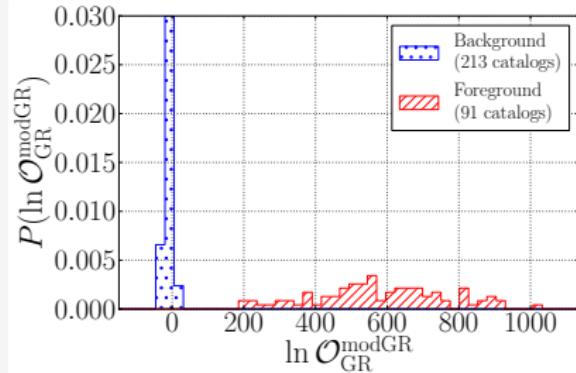
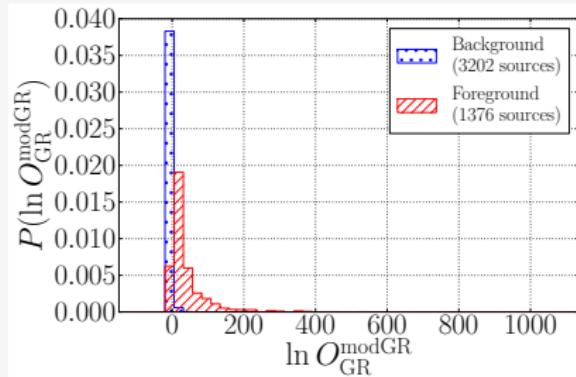
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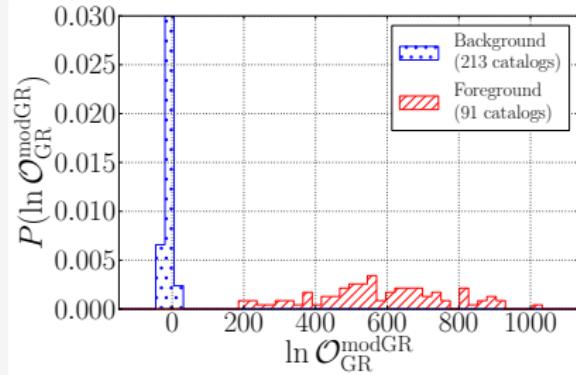
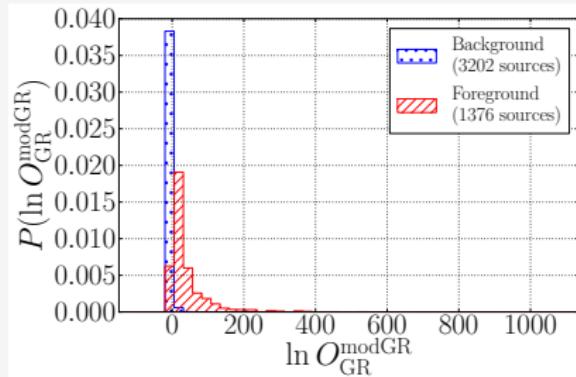
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Other potential issues

Neglected physical effects can cause a perceived violation of GR

- Spin: major contributions expected from 1.5PN
⇒ Include spin in the waveforms
- Tidal distortions: matter effects from 5PN, large pre-factor makes effect significant (BNS/BHNS)
- Merger/ringdown: effects beyond inspiral
⇒ Introduce a frequency cut-off of 400Hz
- Calibration error: mischaracterisation of detector
⇒ Construct the background with calibration errors

Spinning waveforms

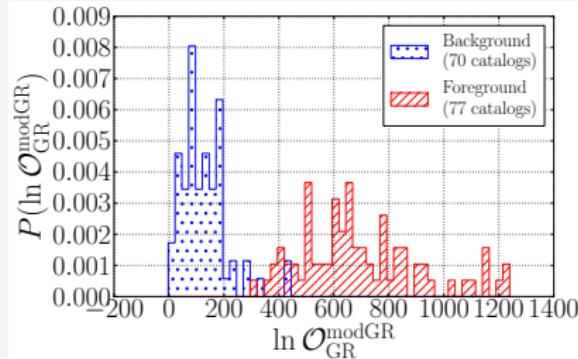
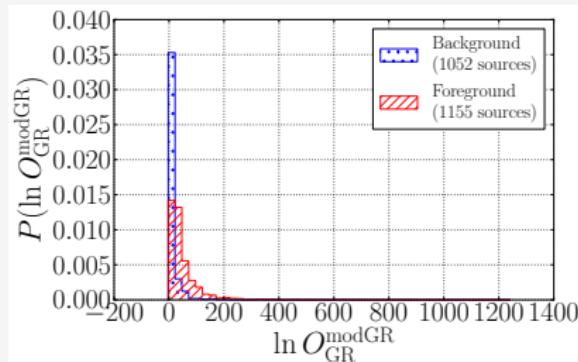
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- Dimensionless spin:
Known spin frequencies,
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 $\mathcal{N}(\mu = 0, \sigma = 0.05)$
- -10% deviation in 1.5PN

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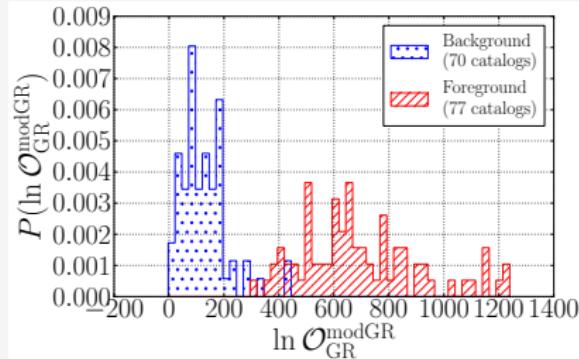
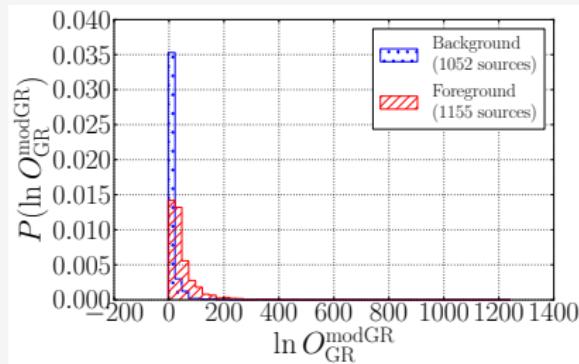
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Upper frequency cut-off

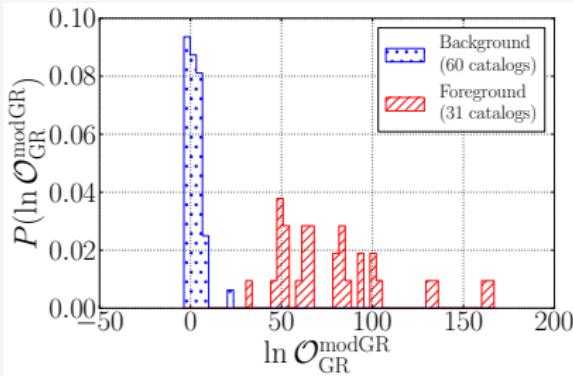
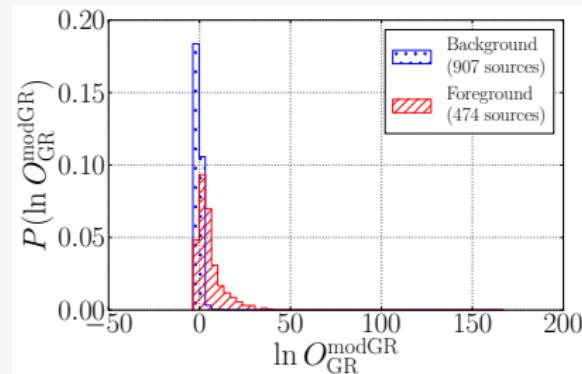
- TaylorF2 templates
- Upper cut-off $f = 400\text{Hz}$
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 - Matter effects [9]
- Bulk SNR from $< 400\text{Hz}$
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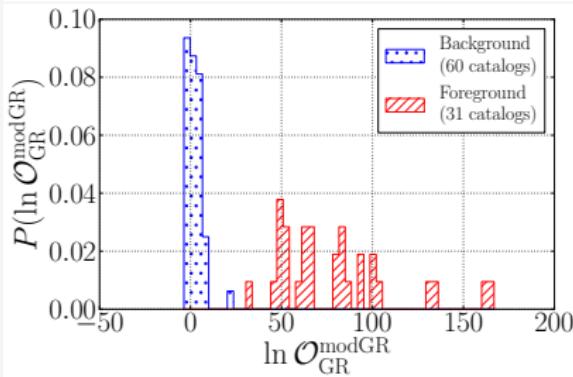
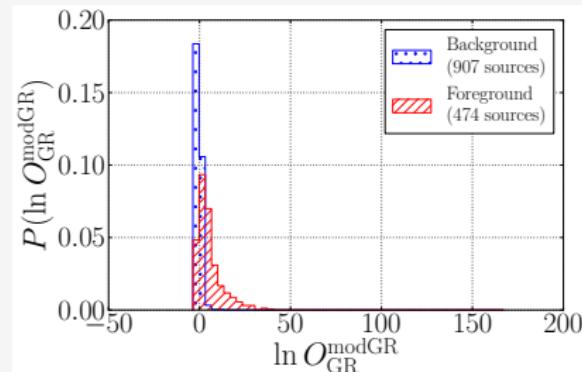
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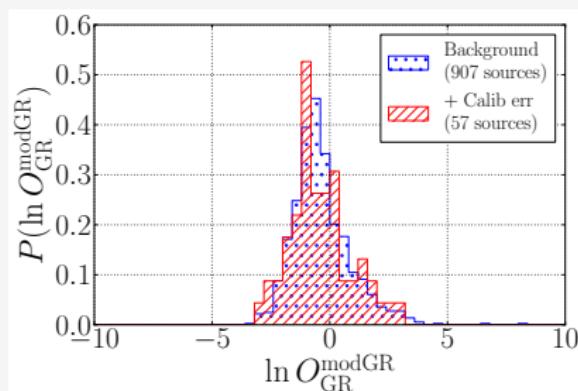
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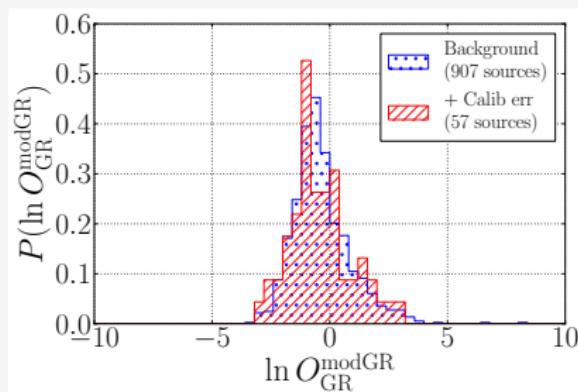
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Merger/ringdown

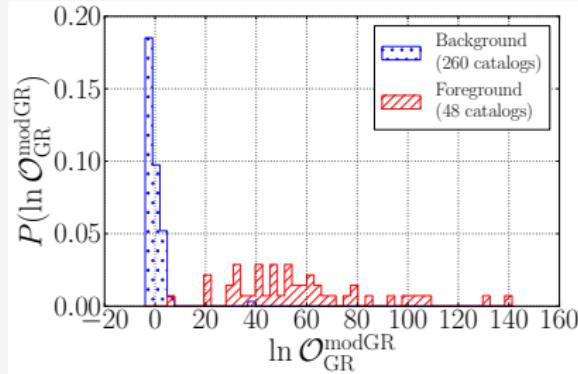
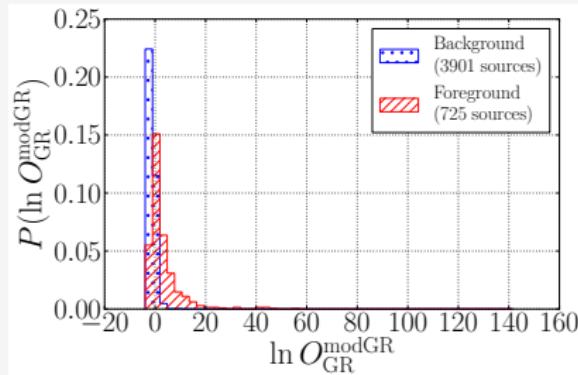
- IMRPhenomB templates [13]
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Merger/ringdown potential
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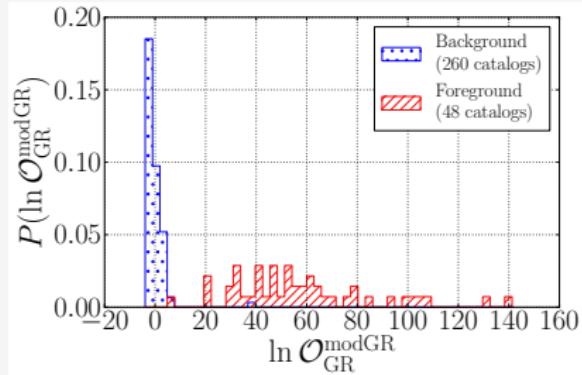
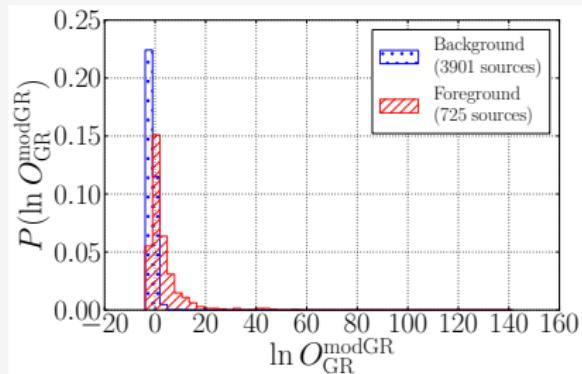
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Conclusions

TIGER: A data analysis pipeline for testing general relativity using compact binary coalescence [1, 2]

- Detect arbitrary deviations
- Suitable for use with any waveform family
- Suitable for Adv LIGO/Virgo (low SNR, moderate amount of sources)
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Join the effort!

Weekly telecons - Thursday
5pm CET / 11am EST / 8am PST